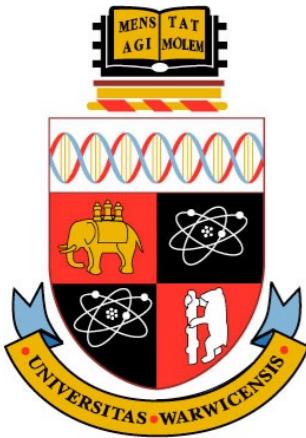


# How PLATO's asteroseismic stellar age constraints could track planetary evolution

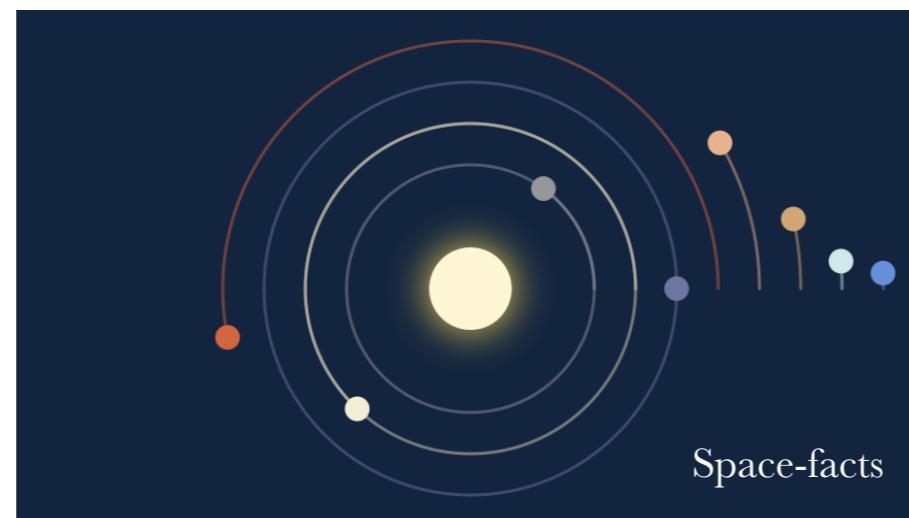


Dimitri Veras  
University of Warwick

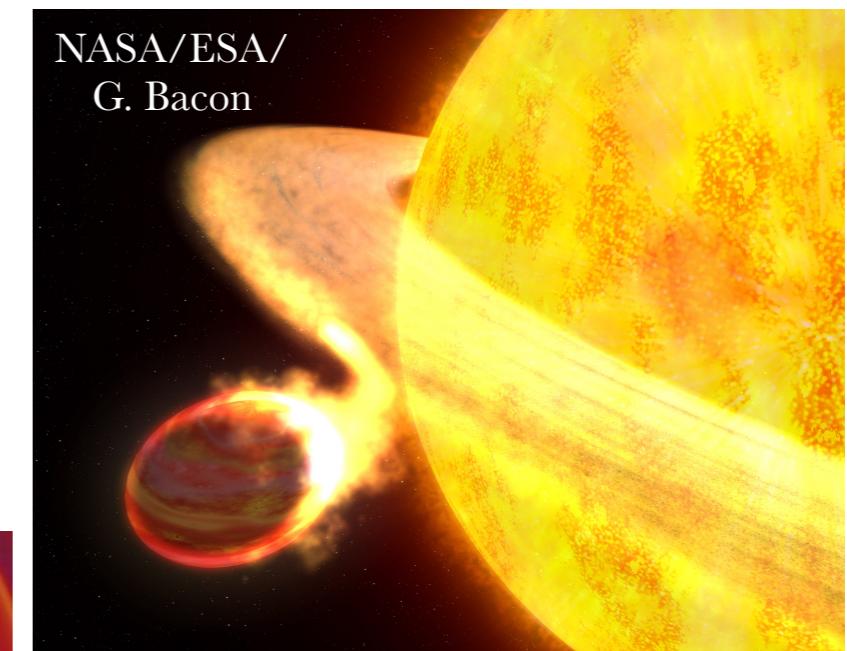


# What can asteroseismology do for planets?

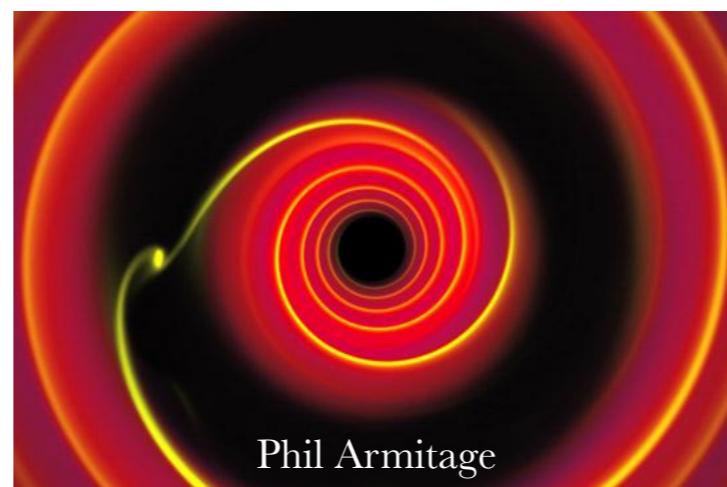
— Tides



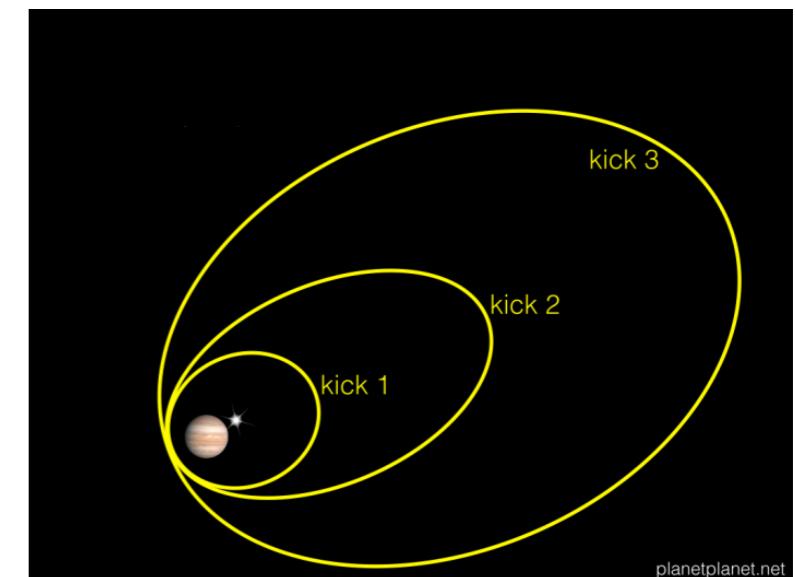
— Planet destruction



— Formation



— Gravitational scattering



# What can PLATO do for asteroseismology?



F5-K7 stars

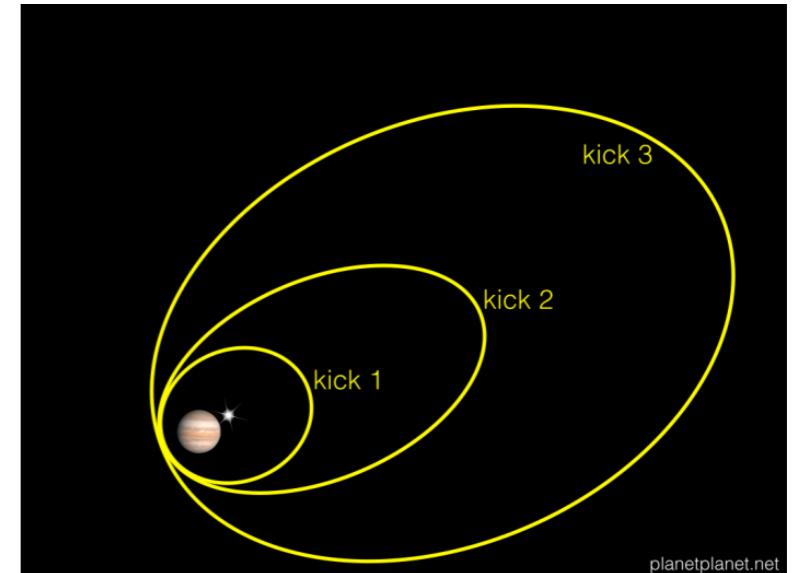
Core sample of ~15,000 stars

Uses *Gaia* radii measurements

10% main sequence age precision

# What can asteroseismology do for planets?

## — Gravitational scattering



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MNRAS **453**, 67–72 (2015)

doi:10.1093/mnras/stv1615

## Prospects for detecting decreasing exoplanet frequency with main-sequence age using *PLATO*

Dimitri Veras,<sup>1</sup>★ David J. A. Brown,<sup>1,2</sup> Alexander J. Mustill<sup>3</sup> and Don Pollacco<sup>1</sup>

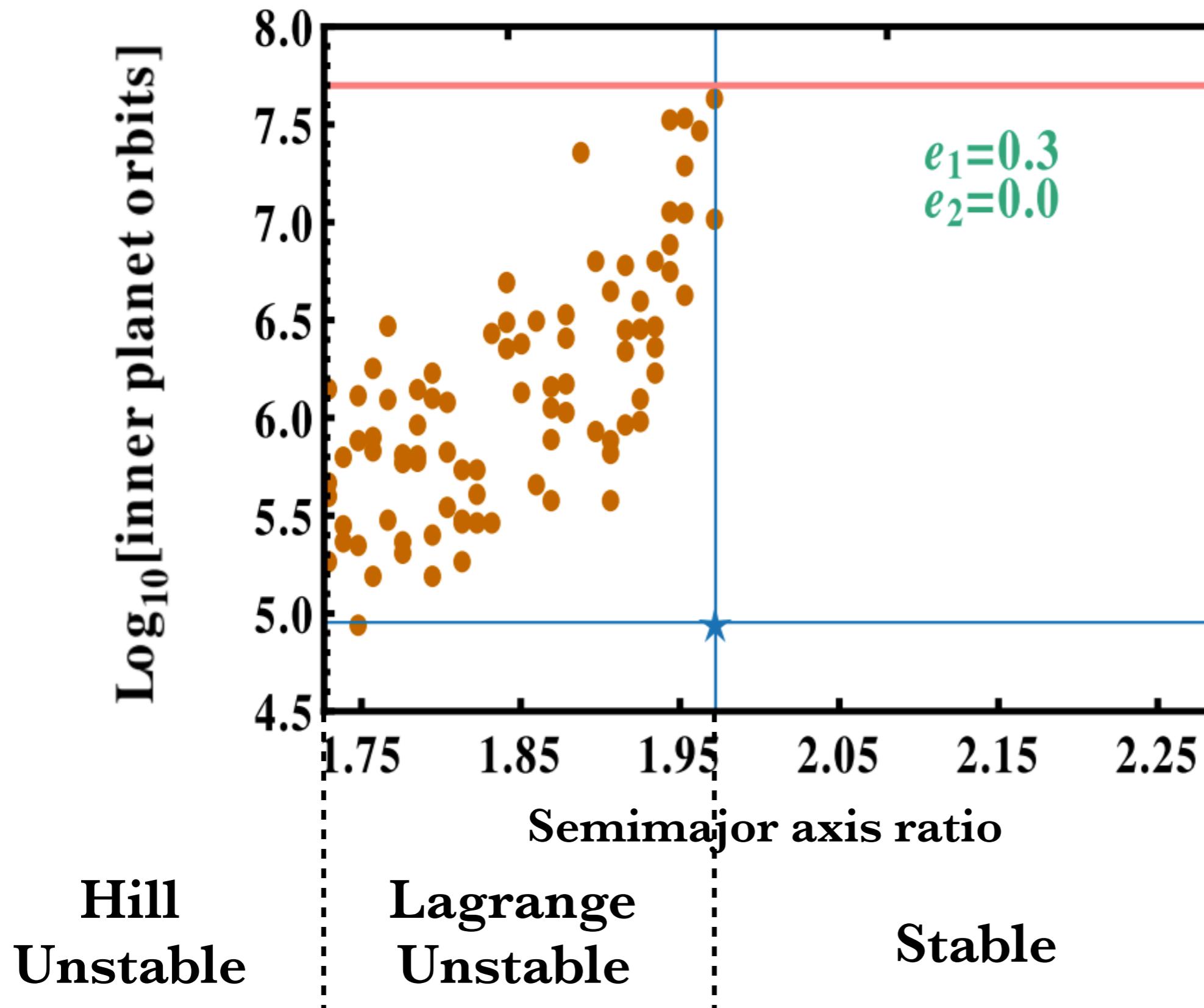
<sup>1</sup>Department of Physics, University of Warwick, Coventry CV4 7AL, UK

<sup>2</sup>Astrophysics Research Centre, School of Mathematics & Physics, Queen's University Belfast, University Road, Belfast BT7 1NN, UK

<sup>3</sup>Lund Observatory, Department of Astronomy and Theoretical Physics, Lund University, Box 43, SE-221 00 Lund, Sweden

# 2-planet unstable systems

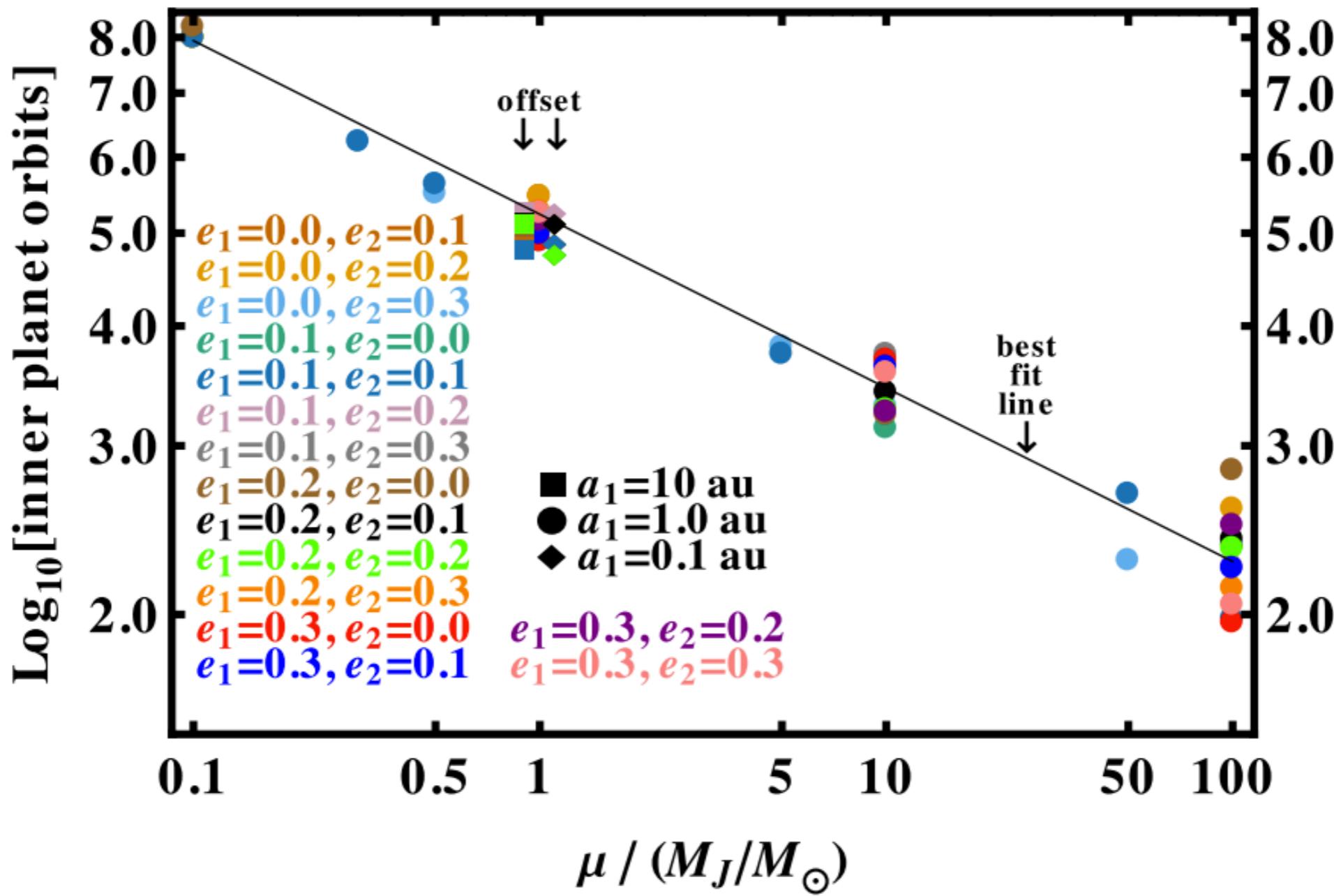
Veras & Mustill (2013, MNRAS, 434, L11)



# 2-planet unstable systems

Veras & Mustill (2013, MNRAS, 434, L11)

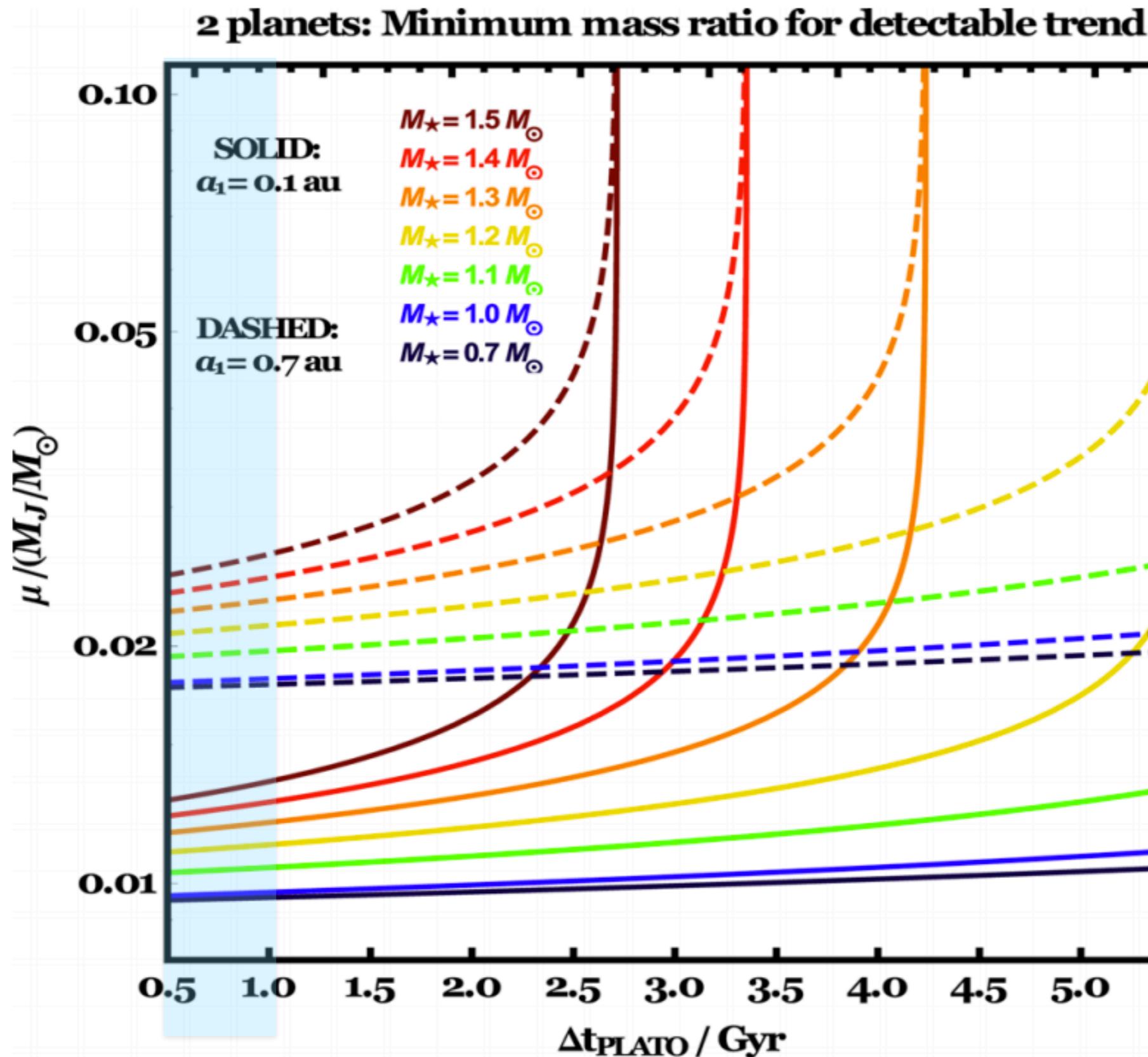
## Minimum Instability Times



$$\log_{10} x \sim 5.2 \left( \frac{\mu}{M_J/M_\odot} \right)^{-0.18}$$

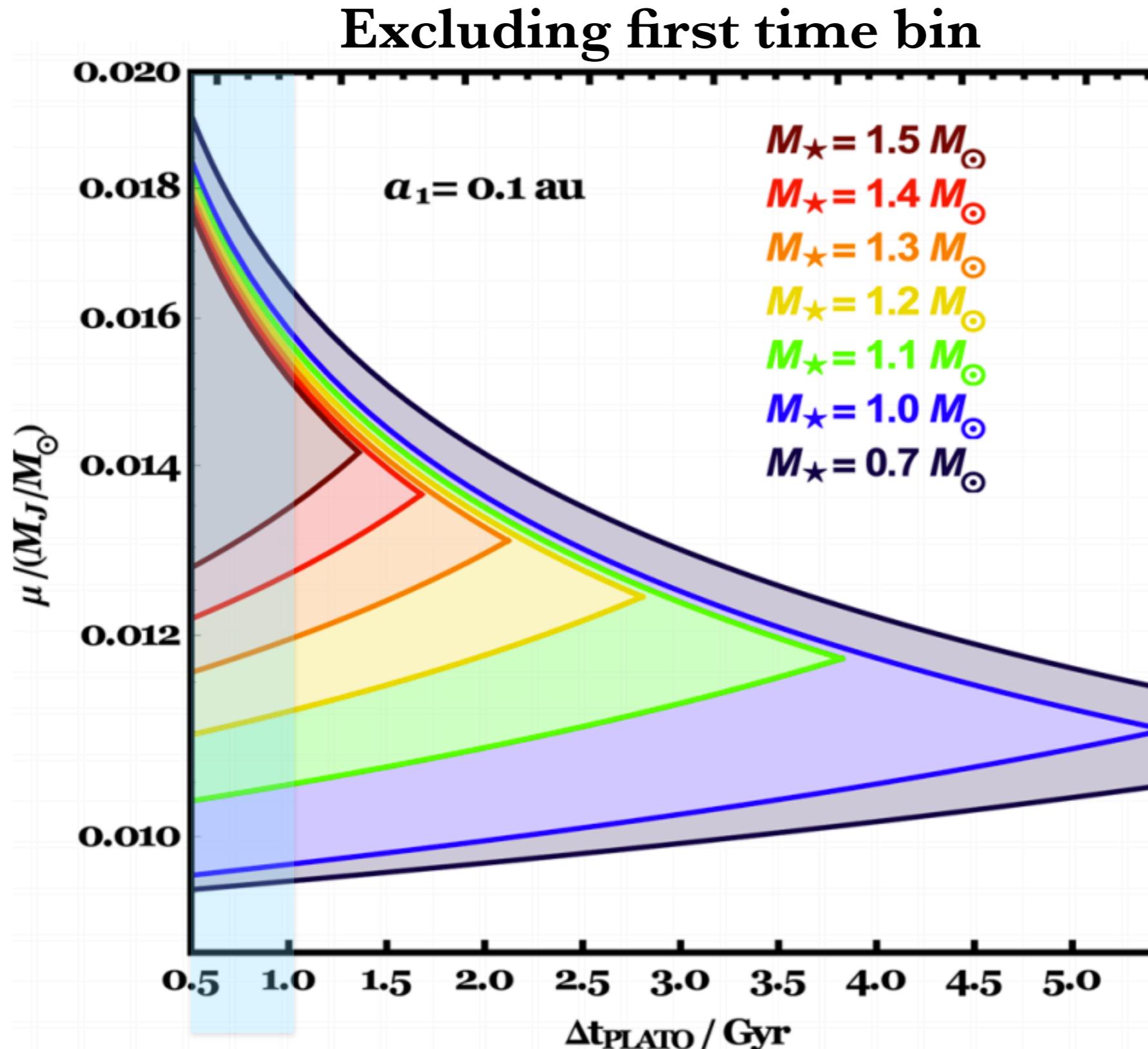
# PLATO asteroseismology with 2 planets

Veras, Brown, Mustill, Pollacco (2015, MNRAS, 453, 67)



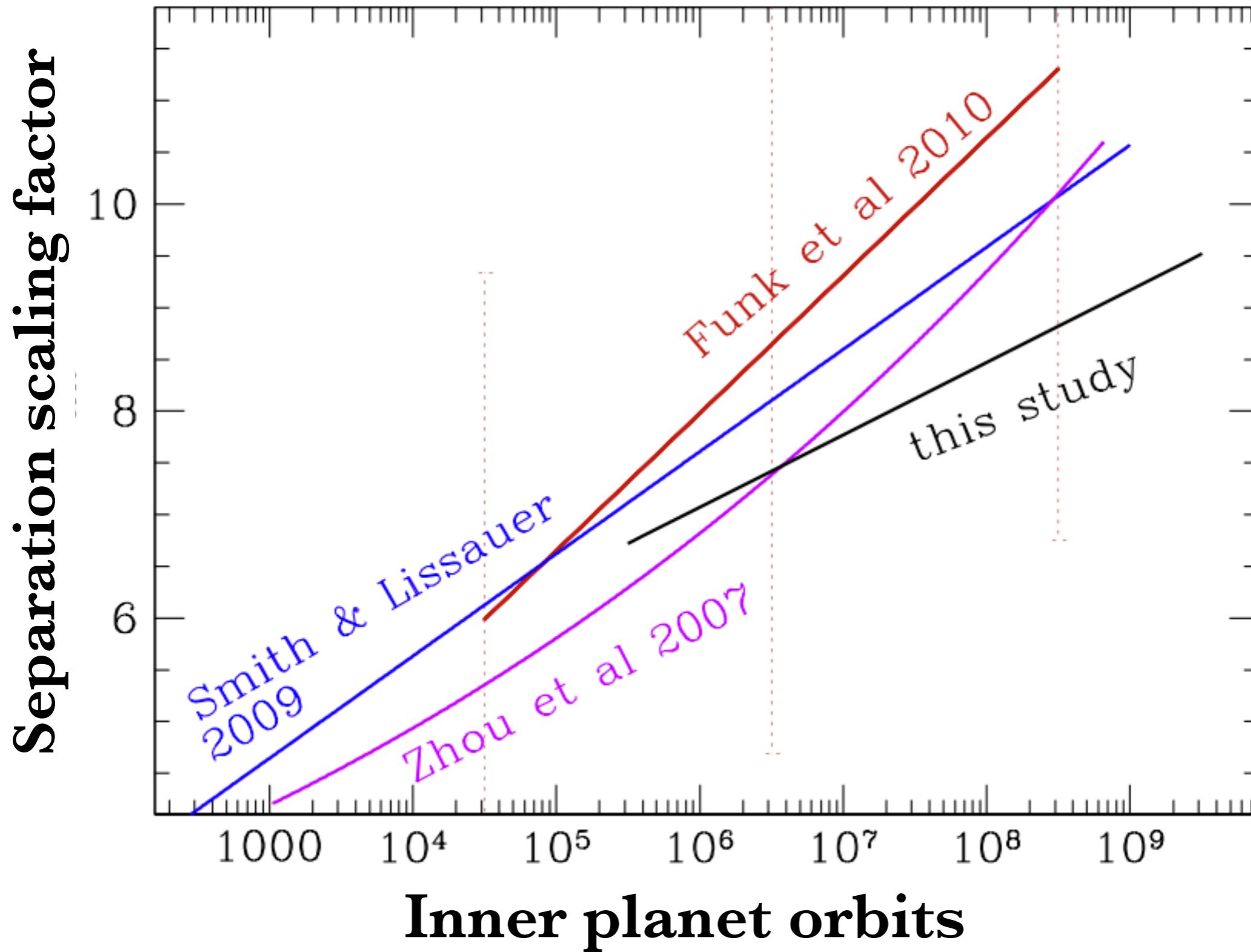
# PLATO asteroseismology with 2 planets

Veras, Brown, Mustill, Pollacco (2015, MNRAS, 453, 67)



# More than 2 planets

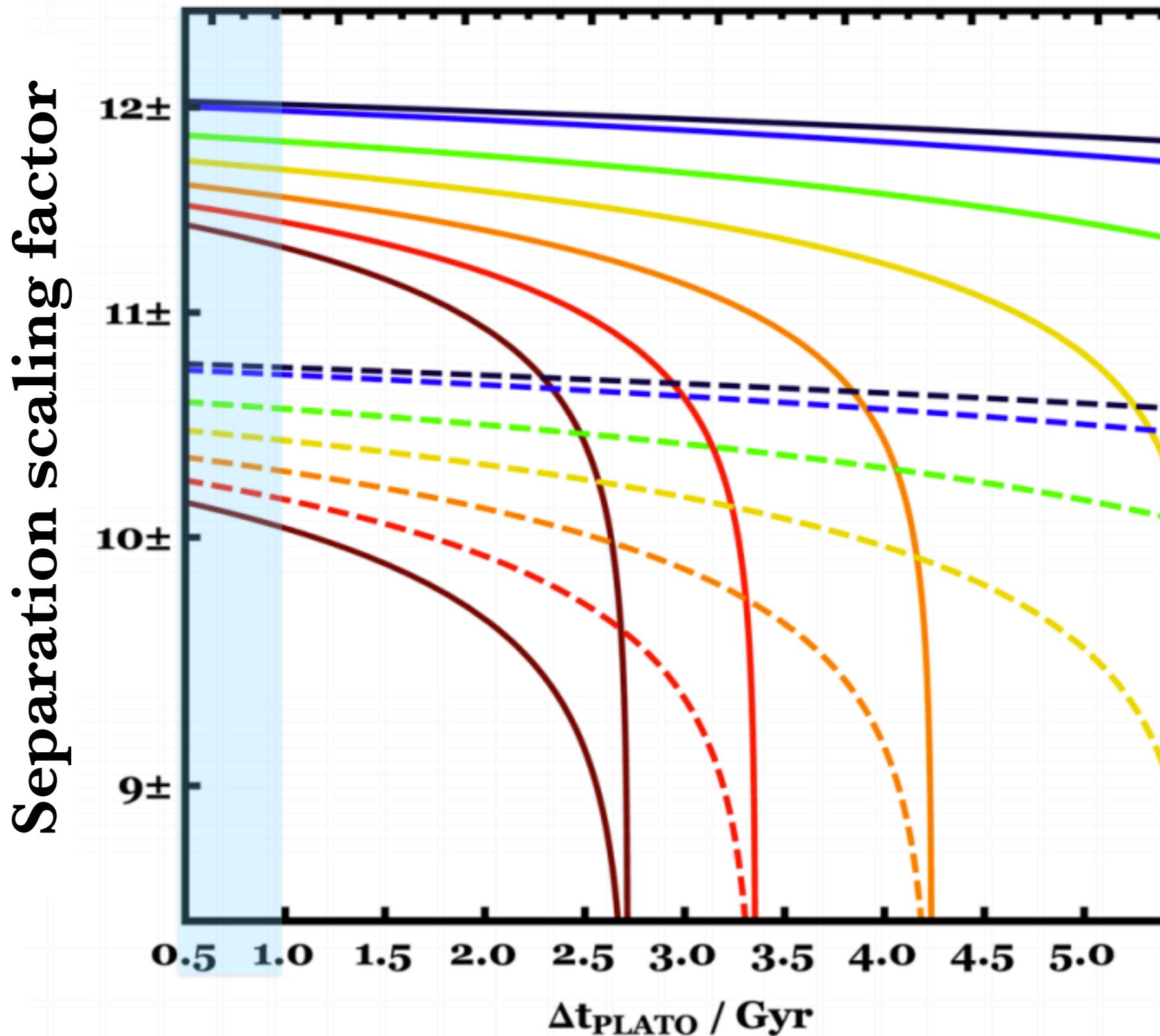
Pu & Wu (2015, ApJ, 807, 44)



# More than 2 planets

Veras, Brown, Mustill, Pollacco (2015, MNRAS, 453, 67)

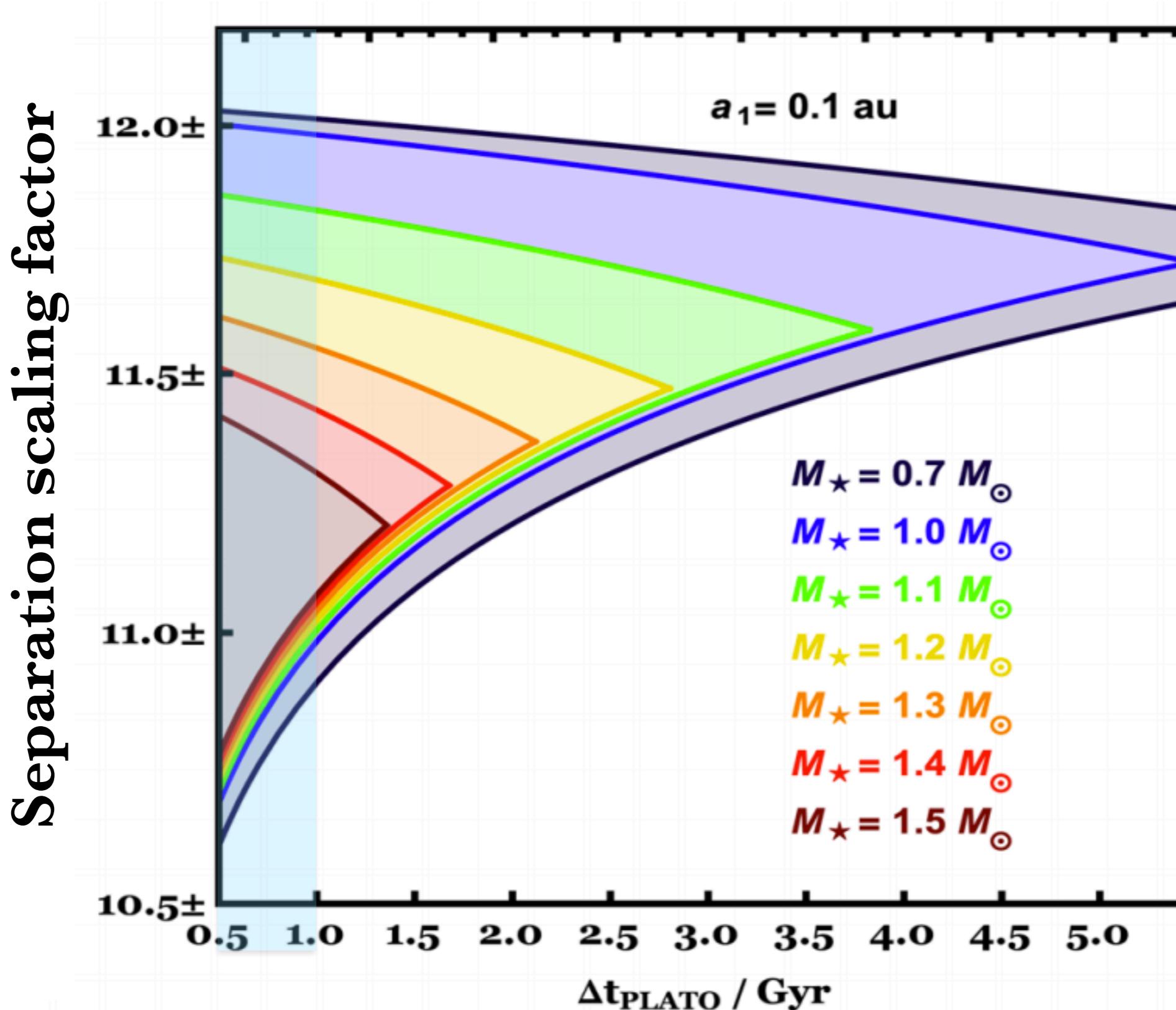
$\geq 3$  planets: Necessary empirical relation for detectable trend



# More than 2 planets

Veras, Brown, Mustill, Pollacco (2015, MNRAS, 453, 67)

Excluding first time bin



# Conclusions

Veras, Brown, Mustill, Pollacco (2015, MNRAS, 453, 67)

**PLATO stellar age constraints can trace planetary system evolution**

**Predict decreasing frequency with time**

**Detectable for ice giants and gas giants**

# What about after main sequence?

9 March 2018  
London

Evolved Solar Systems

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MERGING GIANT-STAR ASTEROSEISMOLOGY WITH THE FATE OF EXTRASOLAR  
PLANETARY SYSTEMS

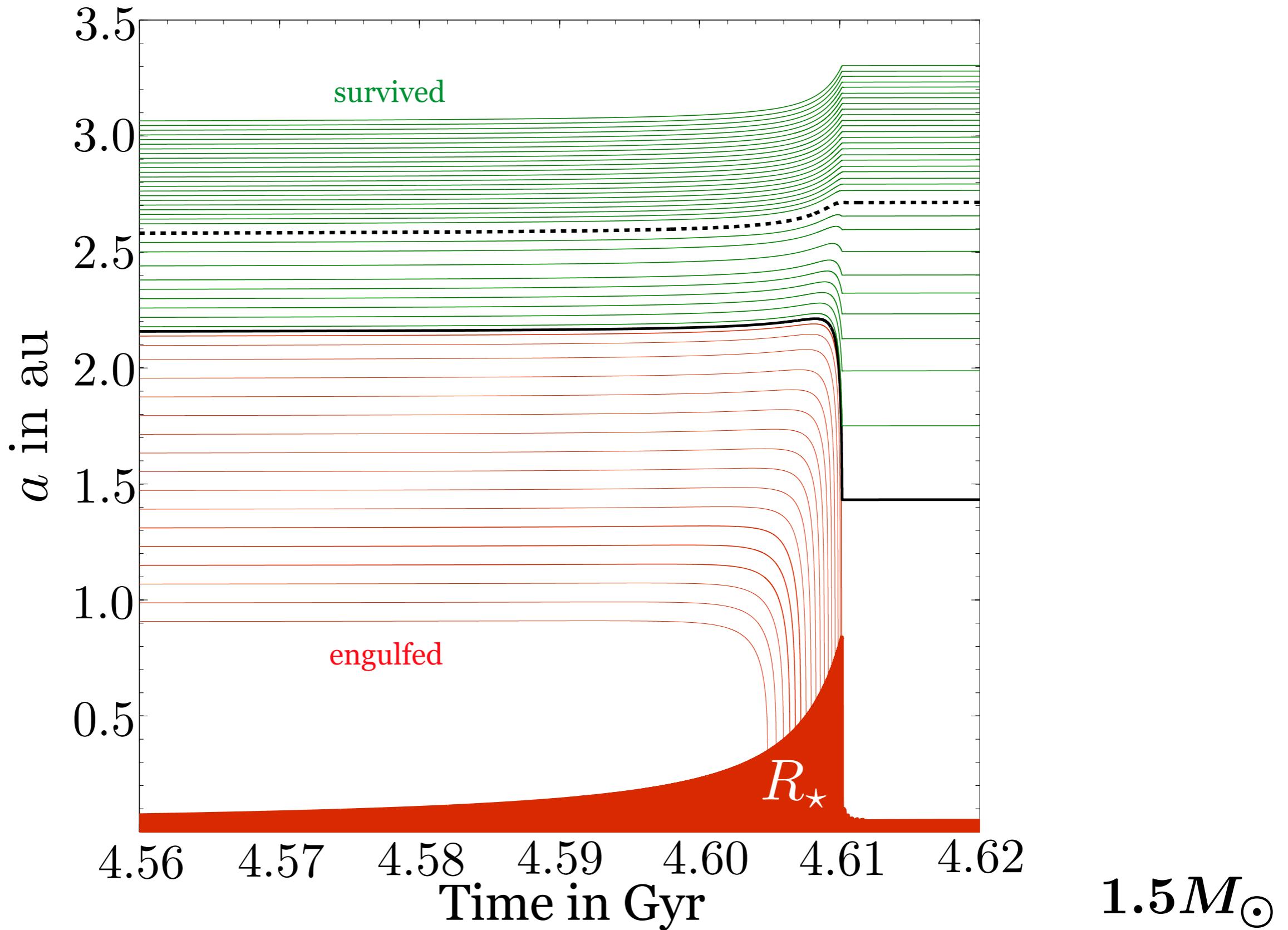
An RAS Specialist Discussion Meeting

*London, UK, March 9th 2018*

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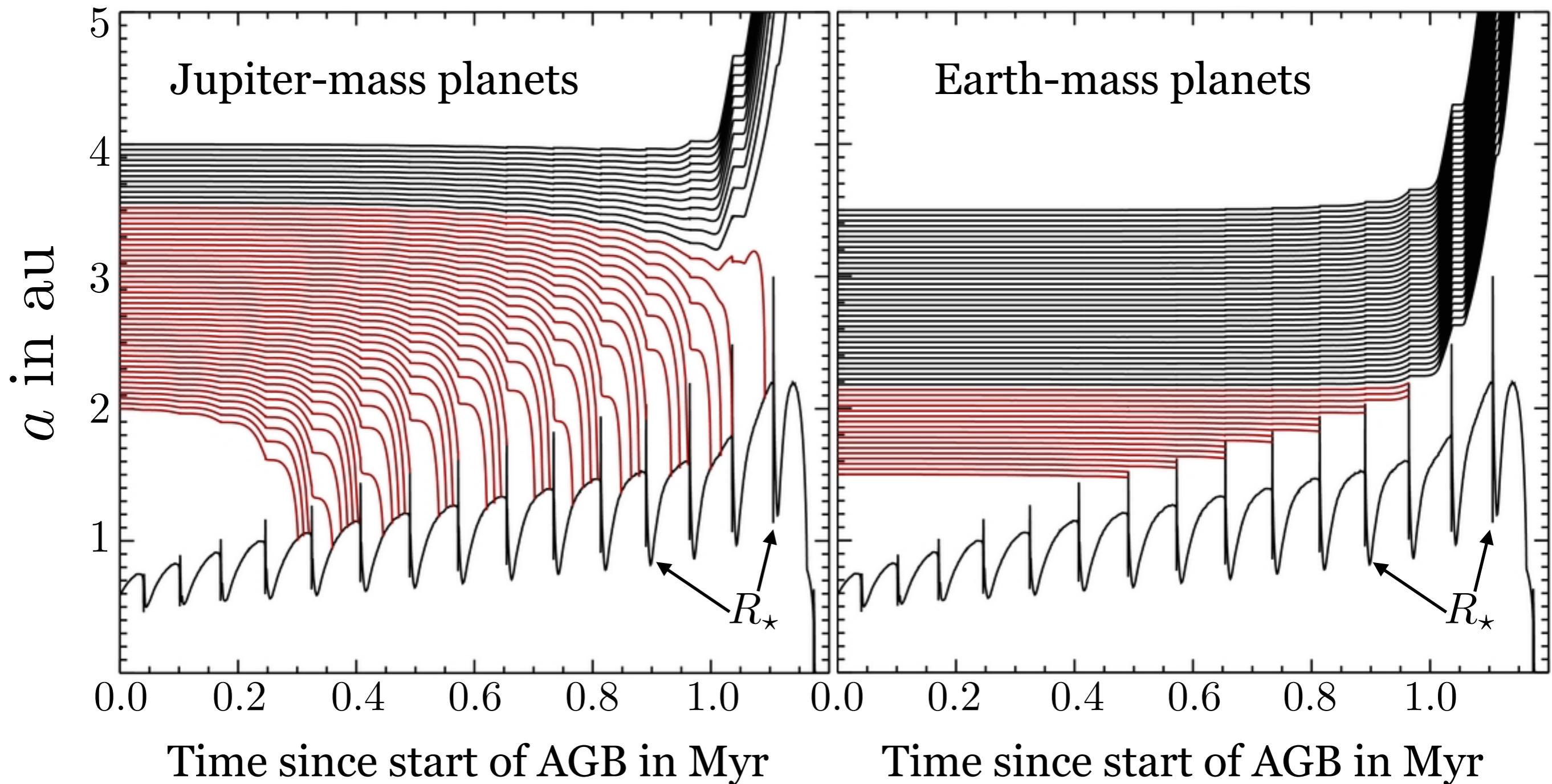
# Fate of planetary systems: red giant branch

Villaver et al. (2014)



# Fate of planetary systems: asymptotic giant branch

Mustill & Villaver (2012)



$2.0 M_{\odot}$