

Limits on radial differential rotation in Sun-like stars from parametric fits to oscillation power spectra

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Martin Bo Nielsen

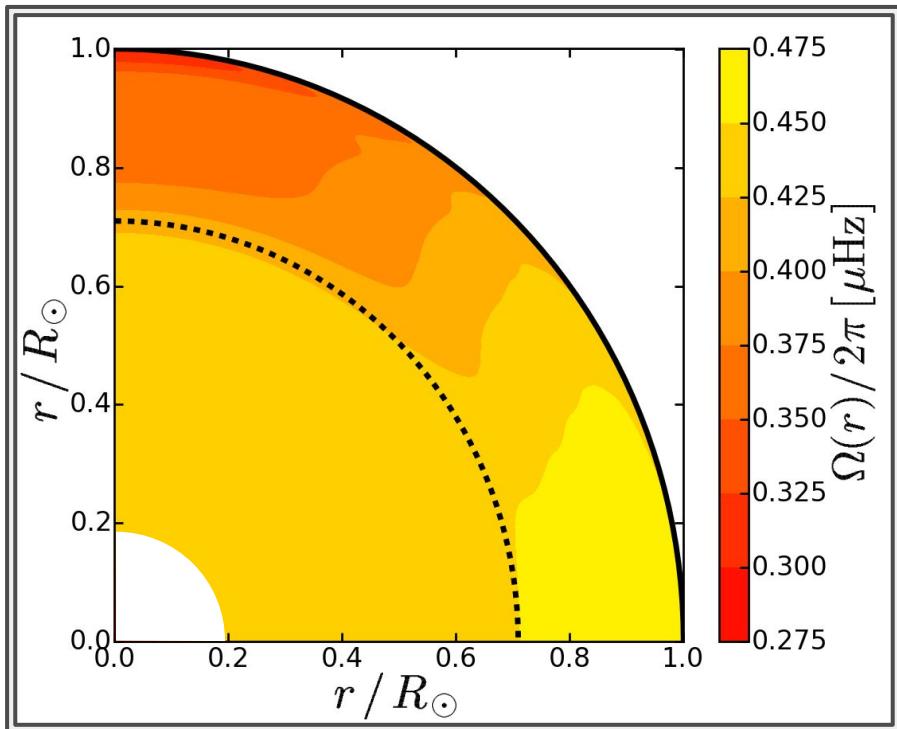
Hannah Schunker, Laurent Gizon, Jesper Schou, Warrick Ball

PLATO mission conference 2017

Rotation in Sun-like stars

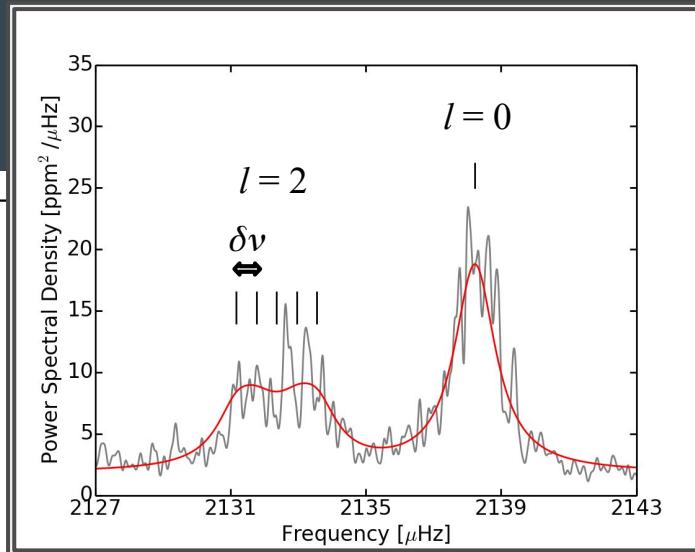
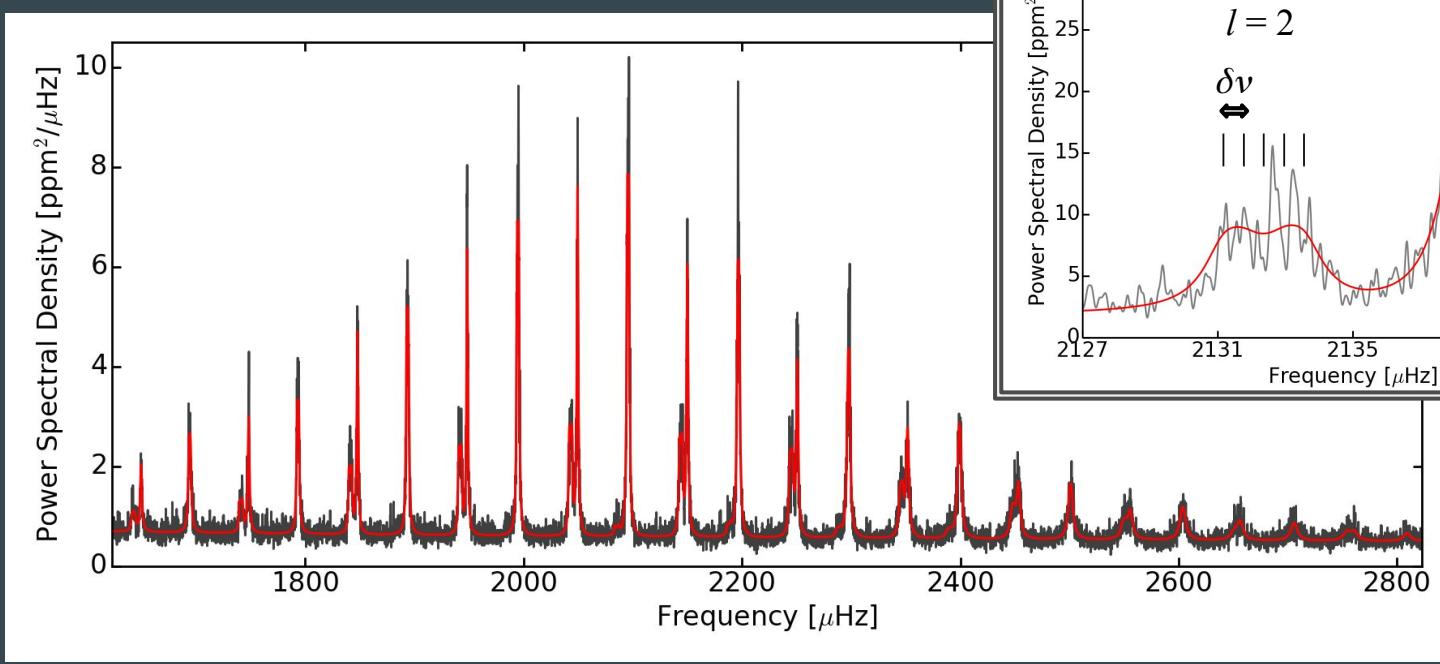
Combine:

- Asteroseismology - sensitive to interior
- Starspot rotation - sensitive to surface

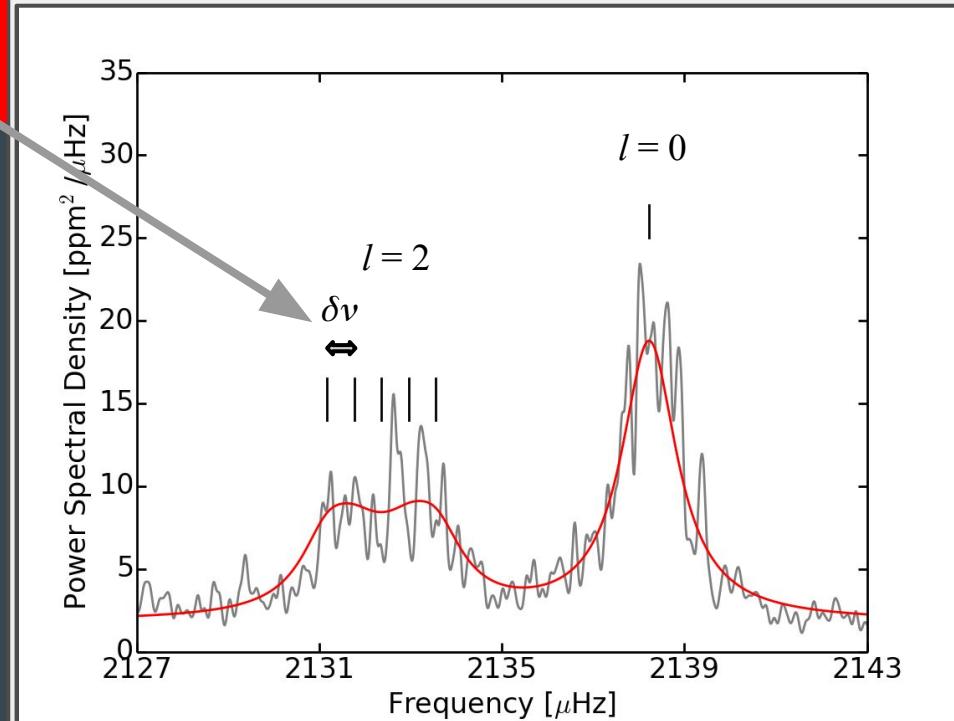
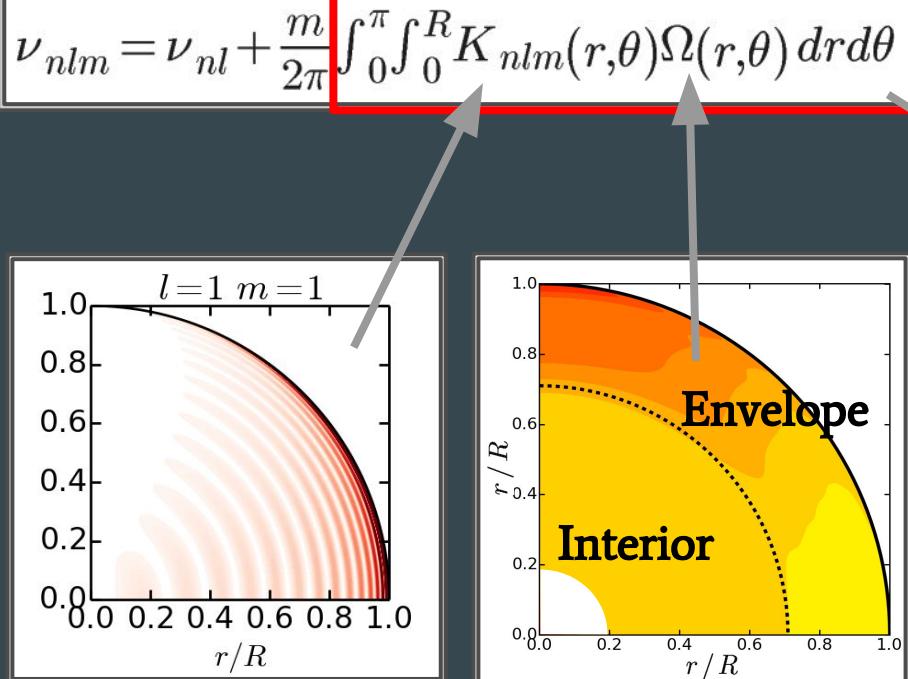


Adapted from Larson & Schou 2015

Measuring rotation with asteroseismology



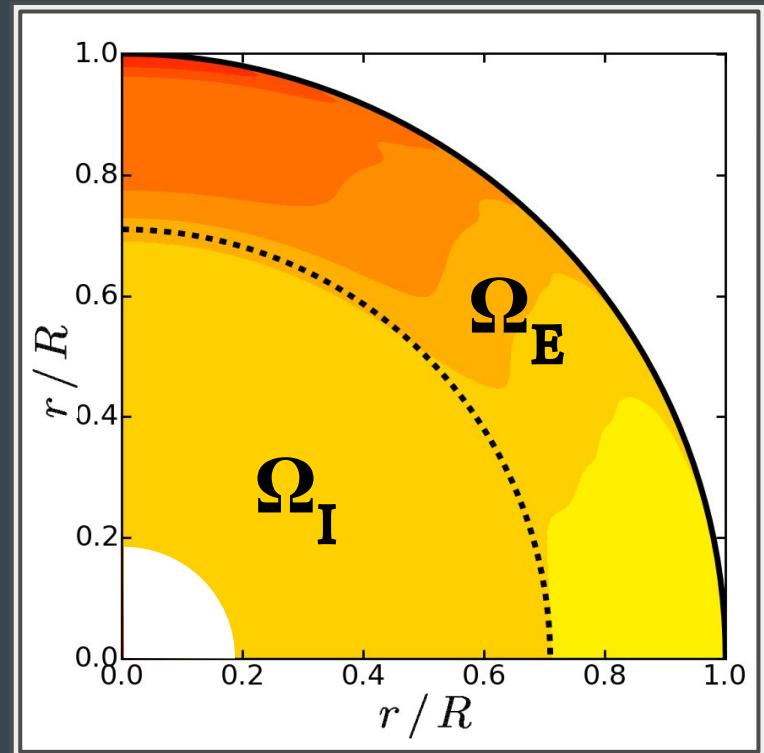
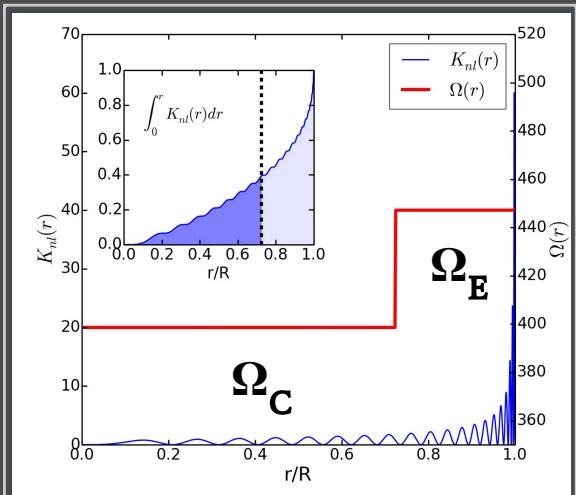
Measuring rotation with asteroseismology



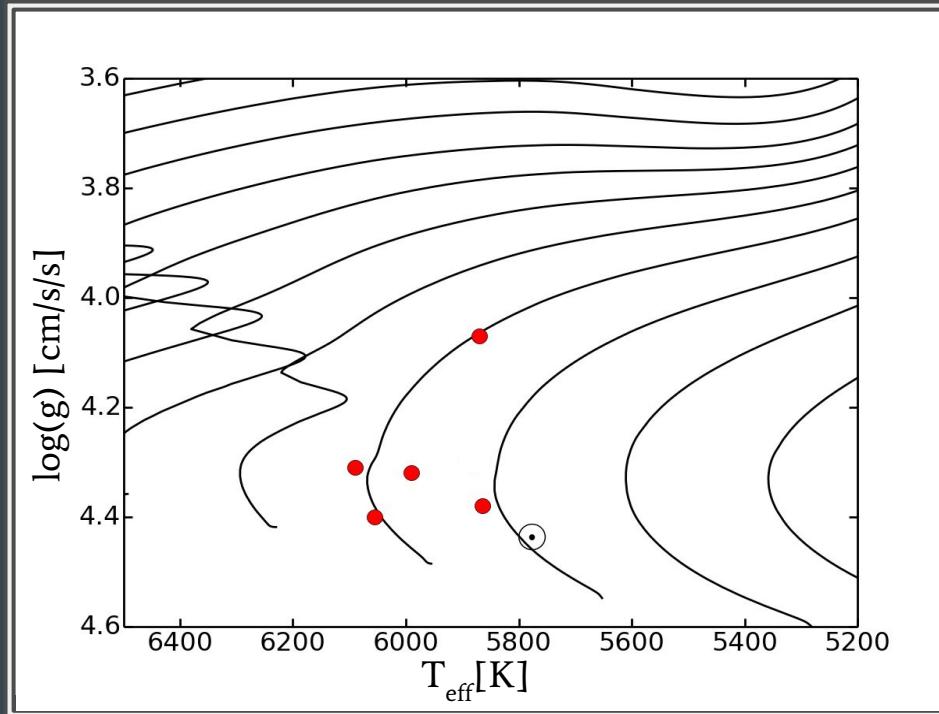
Distinguishing core and envelope rotation

$$\delta\nu_{nlm} = \frac{m}{2\pi} \int_0^R K_{nl}(r) \Omega(r) dr$$

$$\delta\nu_{nlm} = \frac{m}{2\pi} \Omega_I \int_0^{rcz} K_{nl}(r) dr + \frac{m}{2\pi} \Omega_E \int_{rcz}^R K_{nl}(r) dr$$



Selection of Sun-like Kepler stars

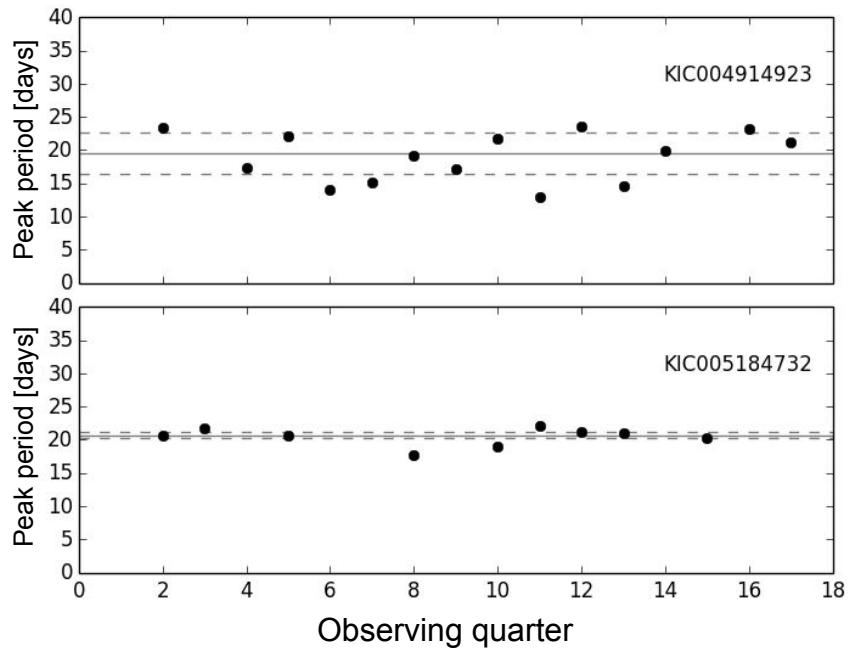
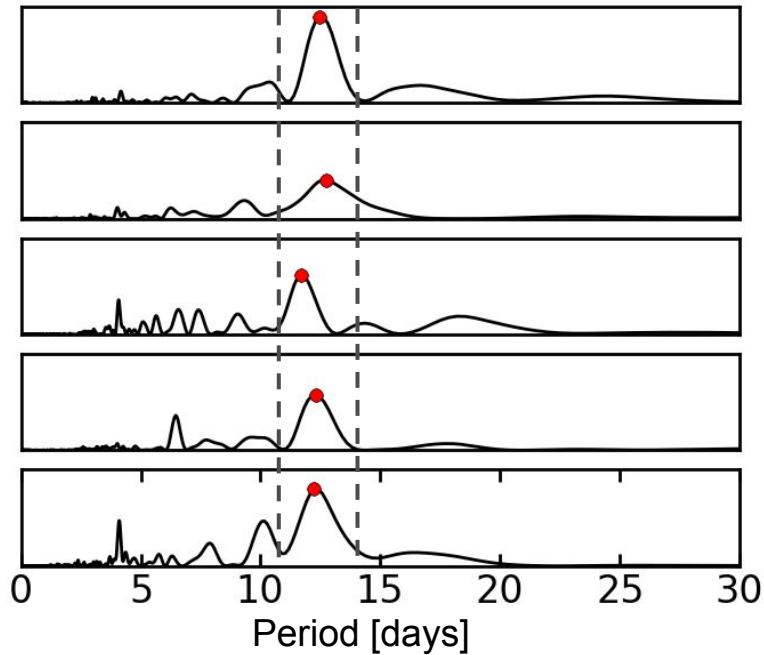


T_{eff} and $\log(g)$ by Bruntt et al. (2012)

- 5 Bright Sun-like stars
- Clear oscillation spectra,
- Clear spot modulation

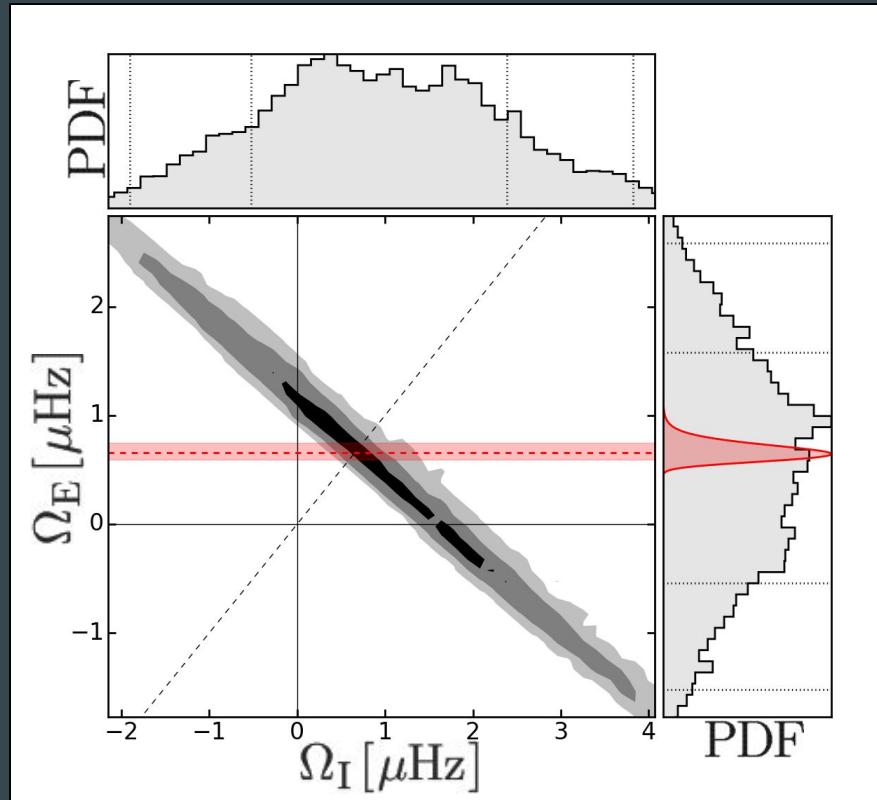
Rotation from surface variability

Power



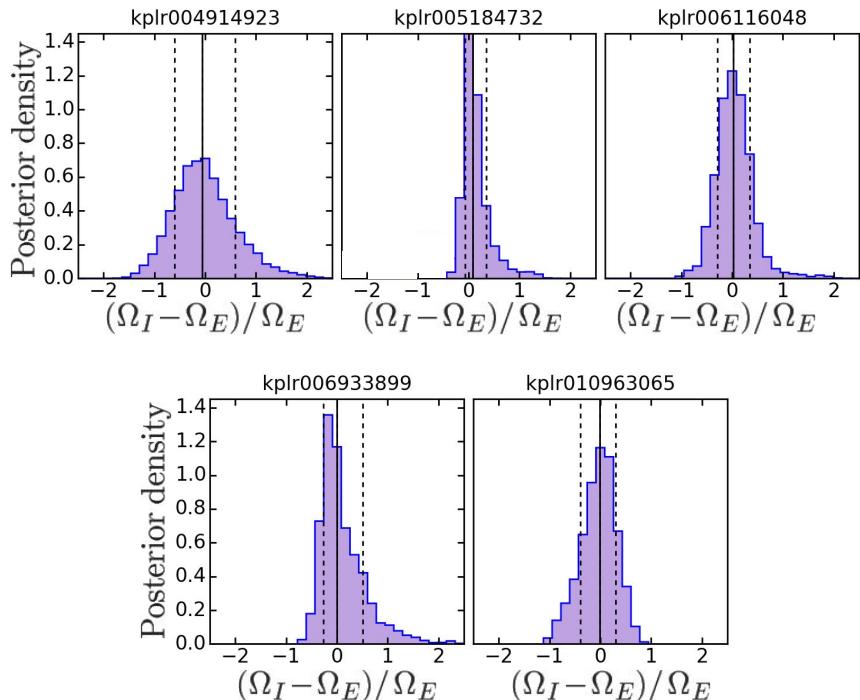
Combining asteroseismology and spots

- Poor constraints from asteroseismology alone
- Surface rotation as prior on Ω_E
- Restricts possible Ω_E and Ω_I



Limits on radial differential rotation

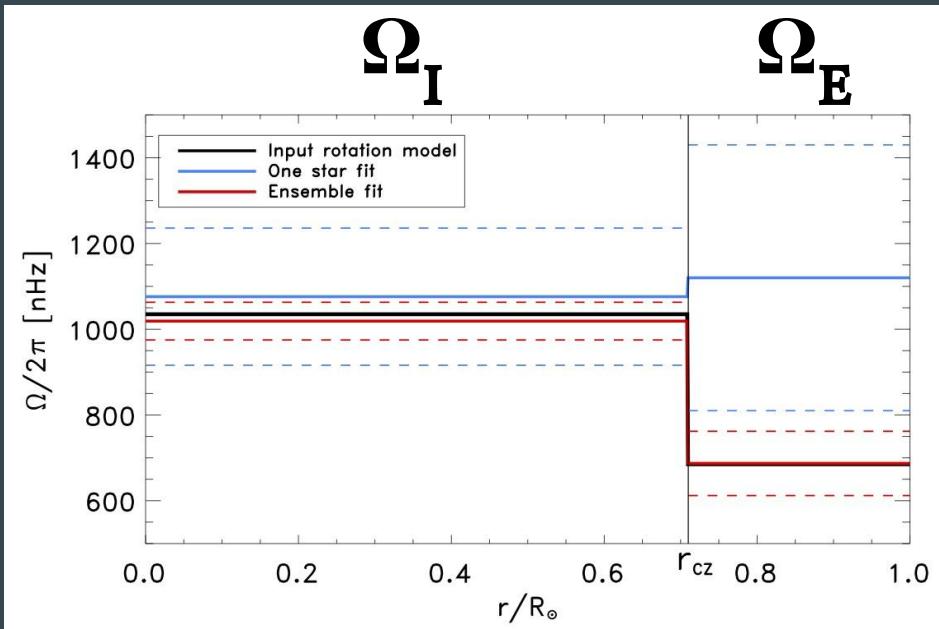
- Centered around 0%
- Average range: $\pm 30\%$
- No clear asymmetry



(Nielsen et al. 2017)

Outlook

- Further reduce shear range
- Ensemble fits (Schunker et al. 2016)
- Simultaneously fit 10-100 stars
- Measure average $\Delta\Omega$ value for all
- PLATO can deliver this



Summary

- ❑ Combined asteroseismology with spot rotation
 - ❑ Shear limited to a range of $\pm 30\%$
 - ❑ PLATO can push this lower for ensembles
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KIC	Cat	Teff [K]	$\log g$ [cm/s/s]
4914923	Vitto	5880 ± 70	4.30 ± 0.08
5184732	Kitty	5865 ± 70	4.38 ± 0.08
6116048	Nunny	5990 ± 70	4.32 ± 0.08
6933899	Fred	5870 ± 70	4.07 ± 0.08
10963065	Rudy	6090 ± 70	4.31 ± 0.08
6106415	Perky	6055 ± 70	4.40 ± 0.08

Comparing seismic and surface rotation rates

- **Average** seismic and surface rotation rates agree within errors
- Constrain the envelope rotation rate in the radial step model

